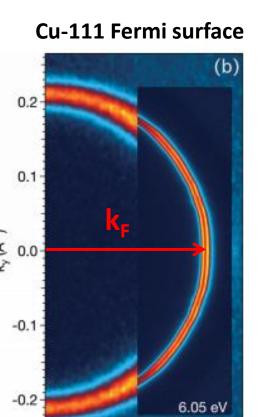
# Angle-resolved photoemission spectroscopy (ARPES)

Overview

#### Outline

Formalism: 3 step Review: model momentum Matrix Looking at space and why elements data: simple we want to go Surface vs bulk there metal Looking at data General principle **ARPES** of ARPES: what we instrumentation do and what we Light source measure Spectrometer Other aspects of Vacuum system experiments Energy/momentum resolution Temperature

# **k** (crystal momentum) vs **q** (momentum transfer)



PRB **87**, 075113 (2013)

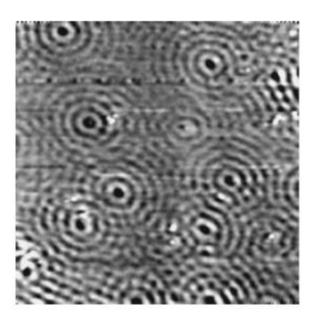
k<sub>x</sub> (Å<sup>-1</sup>)

0.1

0.2

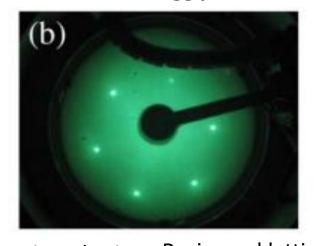
He la

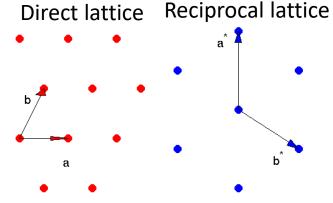
**Cu-111 Friedel Oscillations** 



 $\lambda = \pi/k_F$   $q = 2k_F$ 

**Cu-111 Bragg peaks** 





Thin Solid Films **515** 8285 (2007)

PRB 58 7361 (1998)

### Structures in momentum space

#### All materials

- Brillouin zones
- Fermi surfaces
- Band dispersion

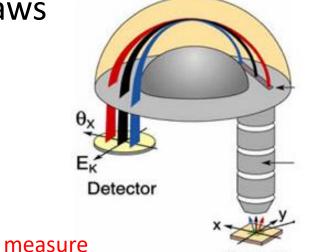
#### Some materials

- Charge density wave gaps (most important for systems without perfect nesting)
- Superconducting gaps
- Spin density wave gaps
- Electron-boson coupling
- Heavy fermion hybridization gaps
- Spin momentum locking
- Surface states
- ...

# Angle-Resolved Photoemission spectroscopy overview

- Purpose: measure electronic band dispersion E vs k
- Photoelectric effect, conservation laws

$$\begin{split} E_{kin} &= h \, \nu - \phi - \frac{|E_B|}{|E_B|} \, |\\ \mathbf{p}_{||} &= \hbar \frac{|\mathbf{k}_{||}}{|E_B|} \cdot \sin \, \mathcal{G} \end{split}$$



want

want

**Definitions:** 

$$E_{kin} = kinetic\ energy\ of\ photoelectron\ measure$$
 
$$hv = photon\ energy\ know$$
 
$$\phi = work\ function\ know/measure$$
 
$$E_B = electron\ binding\ energy\ inside\ material, relative\ to\ Fermi\ level$$
 
$$k_{||} = crystal\ momentum, parallel\ to\ sample\ surface\ plane$$
 
$$m = mass\ of\ free\ electron\ know$$
 
$$\vartheta = emission\ angle\ of\ photoelectron\ measure$$

# What is actually being measured by ARPES?

- Electrons live in bands
- Interactions (electron-electron, electron-phonon, etc) can change band dispersions and quasiparticle lifetimes
- Single particle spectral function captures these interactions

Single particle spectral function: 
$$A(\mathbf{k}, \omega) = -\frac{1}{\pi} \frac{\sum^{"}(\mathbf{k}, \omega)}{\left[\omega - \varepsilon_{\mathbf{k}} - \sum^{'}(\mathbf{k}, \omega)\right]^{2} + \left[\sum^{"}(\mathbf{k}, \omega)\right]^{2}}$$

Bare band:  $\mathcal{E}_{\mathbf{k}}$ 

Self Energy: 
$$\Sigma(\mathbf{k},\omega) = \Sigma'(\mathbf{k},\omega) + i \Sigma''(\mathbf{k},\omega)$$

**Band position** 

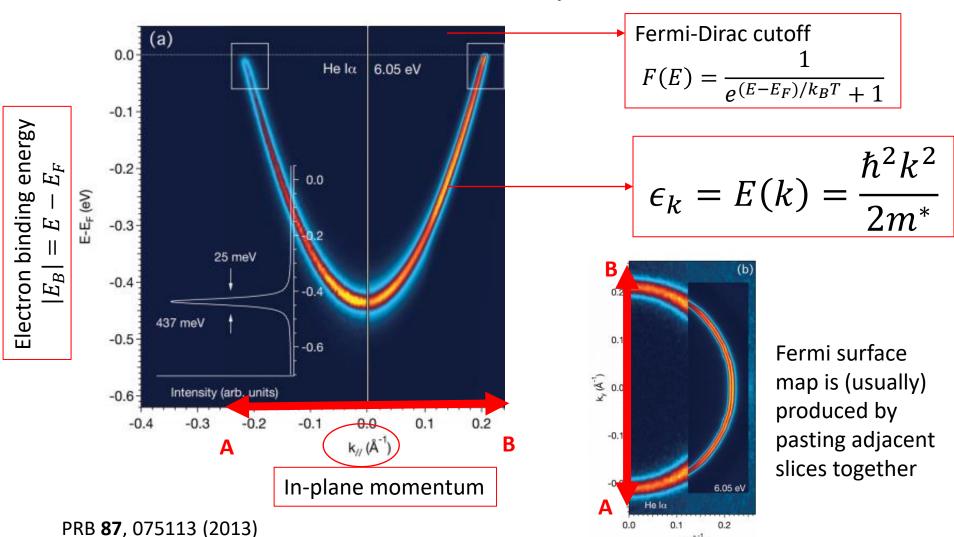
Linewidth or lifetime

Band structure + Interactions

#### Outline

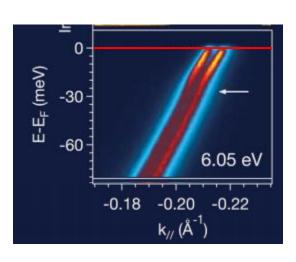
Formalism: 3 step Review: model momentum Matrix Looking at space and why elements we want to go data: simple Surface vs bulk there metal Looking at data General principle **ARPES** of ARPES: what we instrumentation do and what we Light source measure Spectrometer Other aspects of Vacuum system experiments Energy/momentum resolution Temperature

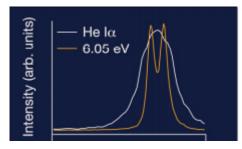
# Band structure: simple metal (Cu 111 surface)



k, (Å-1)

# Self energy: simple metal (Cu 111 surface)



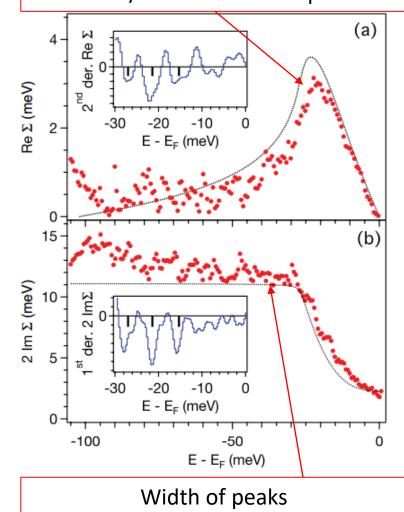


PRB 87, 075113 (2013)

$$A(\mathbf{k}, \omega) = -\frac{1}{\pi} \frac{\sum^{"}(\mathbf{k}, \omega)}{[\omega - \varepsilon_{\mathbf{k}} - \sum^{'}(\mathbf{k}, \omega)]^{2} + [\sum^{"}(\mathbf{k}, \omega)]^{2}}$$

$$\sum (\mathbf{k}, \omega) \rightarrow \sum (\omega) = \sum' (\omega) + i \sum'' (\omega)$$

Measured dispersion minus calculated/assumed bare dispersion

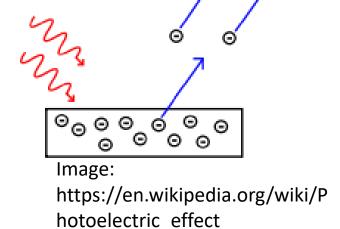


#### Outline

Formalism: 3 step Review: model momentum **Matrix** Looking at space and why elements data: simple we want to go Surface vs bulk there metal Looking at data General principle **ARPES** of ARPES: what we instrumentation do and what we Light source measure Spectrometer Other aspects of Vacuum system experiments Energy/momentum resolution Temperature

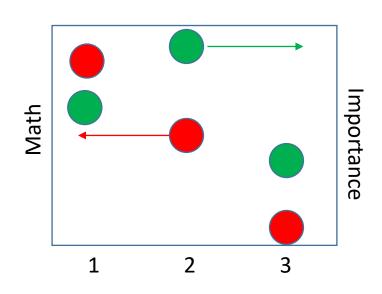
# Back to the beginning: 3 step model

$$\begin{split} E_{kin} &= h \, \nu - \phi - \mid E_B \mid \\ \mathbf{p}_{\parallel} &= \hbar \mathbf{k}_{\parallel} = \sqrt{2mE_{kin}} \cdot \sin \, \mathcal{G} \end{split}$$

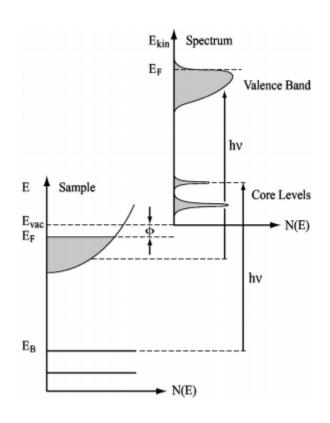


- 1. Optical excitation of electron in the bulk
- 2. Travel of excited electron to the surface
- 3. Escape of photoelectrons into vacuum

Photoemission intensity is given by product of these three processes (and some other stuff)



### 1. Optical excitation of electron in bulk



Hufner. *Photoelectron* Spectroscopy (2003) Start: electron in occupied state of N-electron wavefunction,  $\Psi^N_i$ 

End (of this step): electron in unoccupied state of N electron wavefunction,  $\Psi_f^N$ 

**Sudden Approximation**: no interaction between photoelectron and electron system left behind

Probability of transition related to Fermi's golden rule:

$$w_{fi} = \frac{2\pi}{\hbar} \left| \langle \Psi_f^N \right| - \frac{e}{mc} \mathbf{A} \cdot \mathbf{p} | \Psi_i^N \rangle \right|^2 \delta(E_f^N - E_i^N - h\nu)$$

**p**=electron momentum

**A**=vector potential of photon

Express as product of 1-electron state and N-1 electron state e.g.:  $\Psi_f^N = \mathcal{A}\phi_f^k\Psi_f^{N-1}$ 

# 1. Optical excitation of electron in bulk (continued)

$$\langle \Psi_f^N \left| -\frac{e}{mc} \mathbf{A} \cdot \mathbf{p} \right| \Psi_i^N \rangle = \langle \phi_f^{\mathbf{k}} | -\frac{e}{mc} \mathbf{A} \cdot \mathbf{p} | \phi_i^{\mathbf{k}} \rangle \langle \Psi_m^{N-1} | \Psi_i^{N-1} \rangle$$

$$\equiv M_{f,i}^{\mathbf{k}} \langle \Psi_m^{N-1} | \Psi_i^{N-1} \rangle$$

 $M_{f,i}^{k}$ = 'ARPES matrix elements'=experimental details which affect measured intensity m=index given to N-1-electron excited state with eigenfunction  $\Psi_{m}^{N-1}$  and energy  $E_{m}^{N-1}$ 

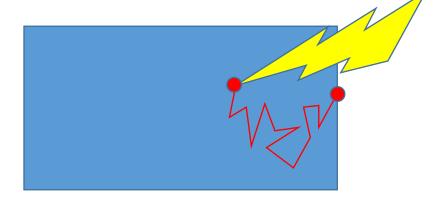
Total photoemission intensity originating from this step:

$$I(\mathbf{k}, E_{kin}) = \sum_{f,i} |M_{f,i}^{\mathbf{k}}|^2 \sum_{m} |\langle \Psi_m^{N-1} | \Psi_i^{N-1} \rangle|^2 \delta(E_{kin} + E_m^{N-1} - E_i^N - h\nu)$$

Consequences of step 1: Observed band intensity is a function of experimental geometry, photon energy, photon polarization

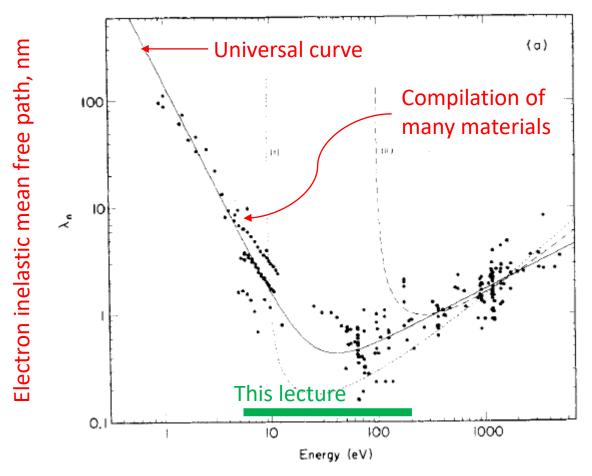
### 2. Travel of excited electron to the

surface



- Excited electrons can scatter traveling to surface
- Typical distance between scattering events = electron mean free path
- What photon energies of light are used in photoemission experiments?
  - 6-6000 eV (this course: 6-150 eV)
- What is the penetration of 20 eV light into copper?
   ~11nm (source: http://xdb.lbl.gov/Section1/Sec\_1-6.pdf)
- What is the electron inelastic mean free path of electrons with kinetic energy 20eV? ~0.6 nm (Seah and Dench)
- What is the size of the Cu unit cell? 0.36 nm

# Electron mean free path universal curve



Seah and Dench, SURFACE AND INTERFACE ANALYSIS, VOL. 1, NO. 1, 1979

Conclusion of Step 2: electron mean free path determines how deep into a sample ARPES studies

**Question:** which photon energy ranges give more bulk sensitivity?

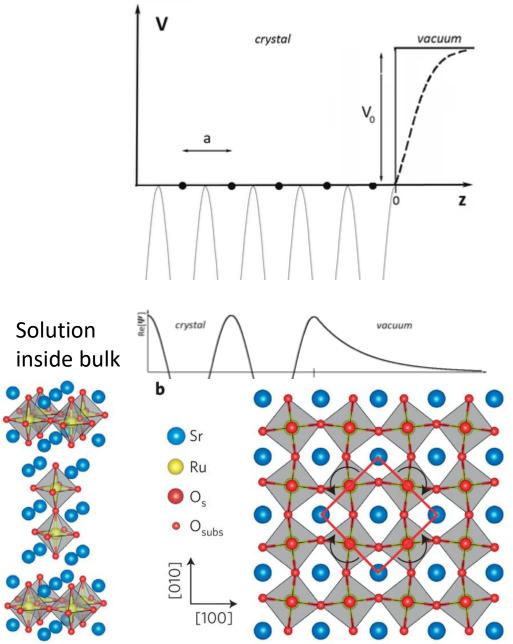
### Surface vs bulk

Inside bulk:  $\Psi_{n,k} = e^{i k \cdot r} u_{n,k}(r)$ 

At surface: deviation from periodicity

#### Various scenarios:

- Electronically distinct state at surface (e.g. Shockley state on Cu 111)
- In quasi-2D materials with weak a coupling between layers, surface termination may not matter much
- Sometimes surface states are interesting (e.g. topological insulators)
- Sometimes atoms on surface will relax/move, changing unit cell



Halwidi et al. Nature Materials 15, 450-455 (2016)

# 3. Escape of photoelectrons into vacuum

- Electron loses work function ( $\Phi$ ) worth of energy
- Transmission probability through surface depends on energy of excited electron and  $\boldsymbol{\Phi}$

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### General setup of ARPES experiment

photon source

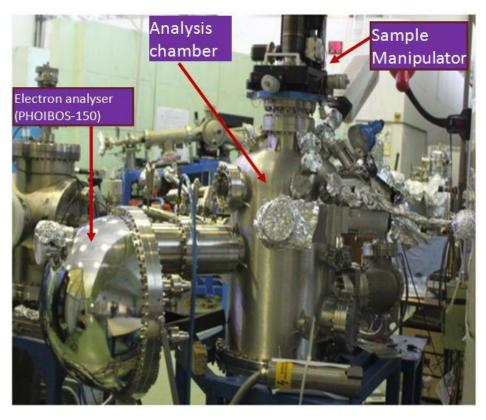


Image source: https://en.wikipedia.org/wiki/Angleresolved\_photoemission\_spectroscopy

кантайн

energy analyses

Image source: http://www.cat.ernet.in/technology/accel/s rul/indus1beamline/arpes.html

## ARPES light sources (6-150 eV)

Type	Available photon energies	Bandwidth/mon ochromaticity	Intensity	Polarization
Laser	6-11 eV; not much variation for a given laser	Can be <<1 meV	Potentially high	Variable polarization
Gas (He, Xe, Ne, Ar) discharge lamp	21.2, 40.8, 8.4, 9.6, 11.6 eV (and more)	Can be small (<1 meV) with monochromator	Sometimes low	random polarization
Synchrotron	Variable; different synchrotrons and endstations specialize in different energy ranges	0.5 to several meV; tradeoff between bandwidth and intensity	tradeoff between bandwidth and intensity	Fixed polarization

$$\begin{split} E_{kin} &= h \, \nu - \phi - \mid E_B \mid \\ \mathbf{p}_{\mid\mid} &= \hbar \mathbf{k}_{\mid\mid} = \sqrt{2mE_{kin}} \cdot \sin \, \mathcal{G} \end{split}$$

$$M_{f,i}^{k} \equiv <\phi_{f}^{k}|-\frac{e}{mc}\boldsymbol{A}\cdot\boldsymbol{p}|\phi_{i}^{k}>$$

### ARPES spectrometer/analyzer



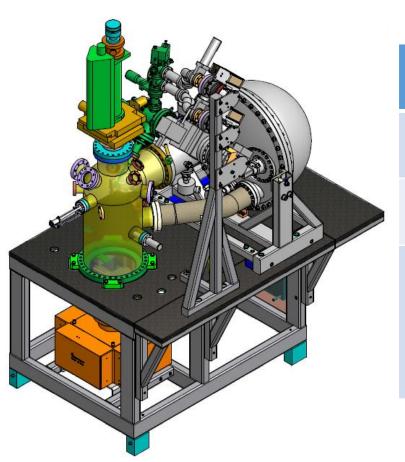


nit.edu/ge rch.html

- Select 1D trajectory in momentum space by rotating sample relative to entrance slit
- Electrostatic lens decelerates and focuses electrons onto entrance slit
- Concentric hemispheres kept at potential difference so that electrons of different energy take different trajectory
- 2D detection of electrons, E vs k

- Electrostatic lens images photoemitted electrons onto position sensitive detector (PSD)
- Discriminate photoelectron energies based on different flight times from sample to detector
- 3D detection of electrons, E vs (kx,ky)

# (Ultra high) vacuum chambers



	High vacuum (HV)	Ultrahigh vacuum (UHV)
Pressure	1e-3 to 1e-9 torr	1e-12 to 1e-9 torr
Molecular mfp	10 cm to 1000km	1000 to 100,000 km
Amount of time to deposit a monolayer on sample surface*	.006s to 95 minutes (typical estimate: 6s)	95 minutes to 65 days (typical estimate: 20 hours)

$$*t = \frac{1.7 \times 10^{-6}}{0.6 \cdot p \cdot S}$$

p=pressure in torr

S=sticking coefficient (between 0 and 1)

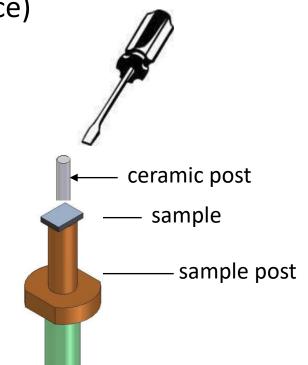
Ref: Hufner, Photoelectron Spectroscopy

### Sample preparation

Achieve atomically clean surface by...

- Cleaving in-situ
- Growing material in-situ
- Sputter-and-anneal (e.g. Cu 111 surface)

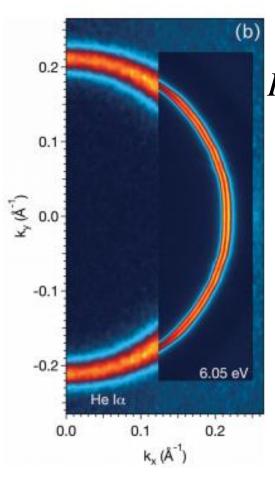
Sample cleaving



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### Resolution in ARPES experiment



Intensity in ARPES experiment:

 $I(\mathbf{k},\omega) = I_0(\mathbf{k},\nu,\mathbf{A}) f(\omega) A(\mathbf{k},\omega) \otimes R(\Delta k, \Delta \omega)$ 

"Matrix elements"

Fermi-Dirac Function Resolution Ellipsoid

Convolution

$$A(\mathbf{k}, \omega) = -\frac{1}{\pi} \frac{\sum^{"}(\mathbf{k}, \omega)}{[\omega - \varepsilon_{\mathbf{k}} - \sum^{'}(\mathbf{k}, \omega)]^{2} + [\sum^{"}(\mathbf{k}, \omega)]^{2}}$$

"band structure + Interactions"

PRB **87**, 075113 (2013)

### Energy resolution

#### Origins of energy broadening

- Light source bandwidth
- Electrical noise
- Spectrometer

$$E_{pass} = \frac{e\Delta V}{\frac{R_1}{R_2} - \frac{R_2}{R_1}} = 0.5, 1, 2, 5, 10 \text{eV}, \text{ or more}$$

$$\Delta E_a = E_{pass} \left( \frac{w}{R_0} + \frac{\alpha^2}{4} \right)$$

w = width of entrance slit (as small as .05 mm)

 $R_0$ =average radius of analyzer (~20 cm)

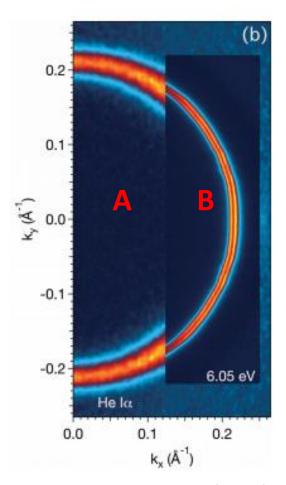
 $\alpha$  =angular resolution (as small as .05°)

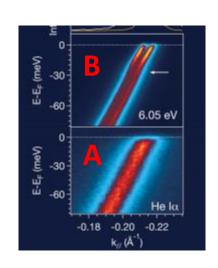
#### Momentum resolution

$$\begin{split} E_{kin} &= h\,\nu - \phi - \mid E_B \mid \\ \mathbf{p}_{\mid\mid} &= \hbar \mathbf{k}_{\mid\mid} = \sqrt{2mE_{kin}} \cdot \sin \mathcal{G} \\ \Delta \mathbf{k}_{\mid\mid} &= \frac{\sqrt{2mE_{kin}} \cdot \cos \mathcal{G}}{\hbar} \,\Delta \mathcal{G} \\ \end{split}$$
 Related to angular resolution of spectrometer and beam spot size

For a given spectrometer, how can one improve momentum resolution?

# Cu 111 ARPES: origin of superior resolution?





Why is B sharper than A?

PRB **87**, 075113 (2013)

#### Some notes on resolution...

- Instrument resolution represents a convolution of original spectrum with 2D resolution ellipsoid. It does not represent the smallest energy or momentum scale which can be resolved
- Resolution can move spectral features around a bit
- There are sometimes tradeoffs to achieving better resolution (e.g. sacrificing photon intensity or ability to access all of momentum space) which may be unacceptable for some experiments
- Resolution has improved a lot in the last 30 years

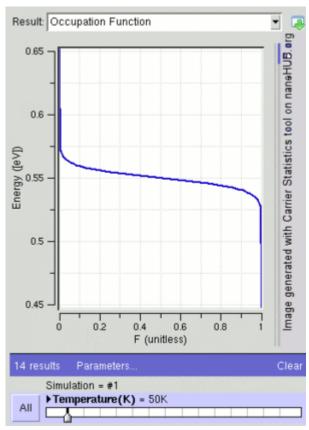
### What about temperature?

$$I(\mathbf{k},\omega) = I_0(\mathbf{k},\nu,\mathbf{A}) f(\omega) A(\mathbf{k},\omega) \otimes R(\Delta k,\Delta \omega)$$

- Fermi-Dirac cutoff gets broader giving access to more unoccupied states
- Spectra get broader, generally following electron lifetime of material system

#### Temperature control during experiment:

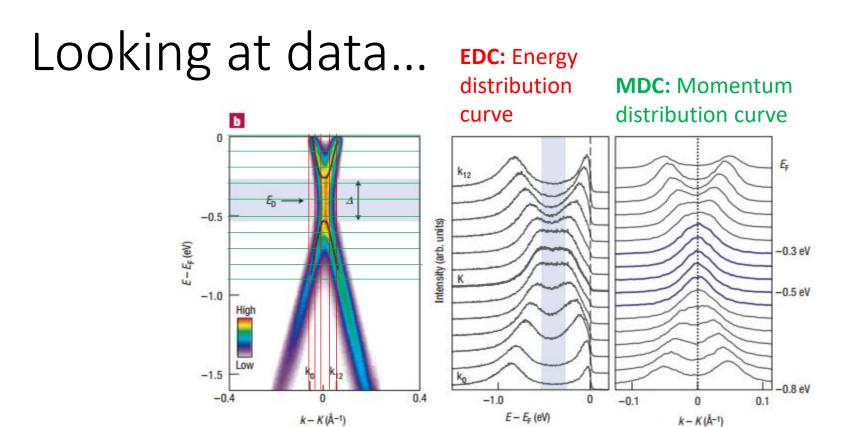
- Flow cryostat
- Maximum temperature ~400K
- Minimum temperature
  - 20K standard
  - ~7K with radiation shielding
  - ~1K high end



Source:

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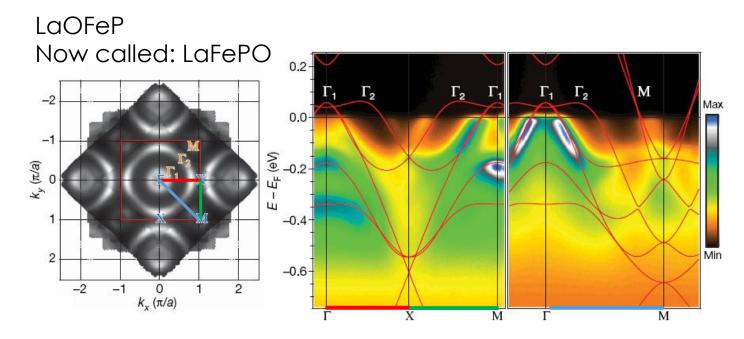


Zhou *et al* Nat. Mater **6** 770 (2007)

Main result: substrate (SiC) breaks sublattice symmetry of graphene, opening a gap at the Dirac point

Which analysis (EDC or MDC) illustrates this result better?

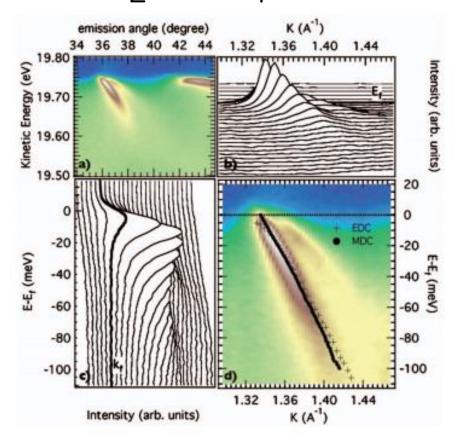
### Looking at more data...



D. H. Lu, et al. Nature **455** 81 (2008)

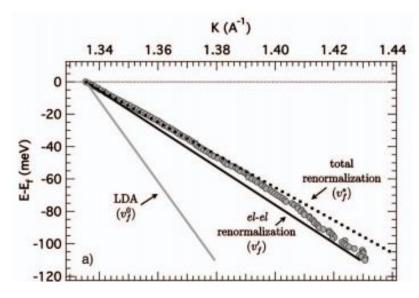
- Data taken along 1D trajectories in k-space (cuts); high-symmetry cuts in these data, but not always
- Fermi surface map produced by pasting many 1D cuts together
- Matrix elements: same band has different brightness for different experiment geometries
- Interaction between experiment and theory

# More data: quantitative analysis of Sr<sub>2</sub>RuO<sub>4</sub> lineshape

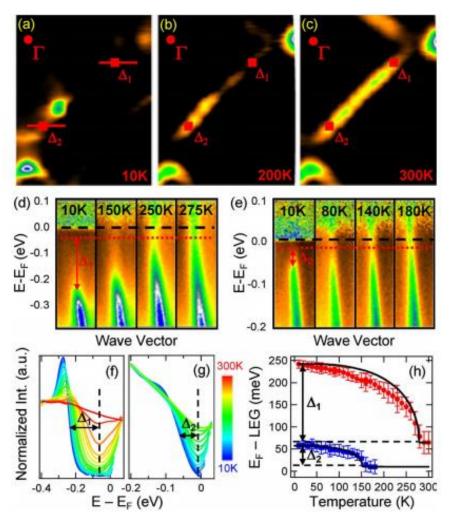


N. Ingle et al. PRB 72, 205114 2005

Why does EDC and MDC analysis give different band position?



# Spectral gaps in ARPES: example ErTe3 (CDW system)



Method 1: Disappearance of intensity in fermi surface map (only works if gap opens at  $E_{\epsilon}$ )

Method 2: Energy vs k dispersion

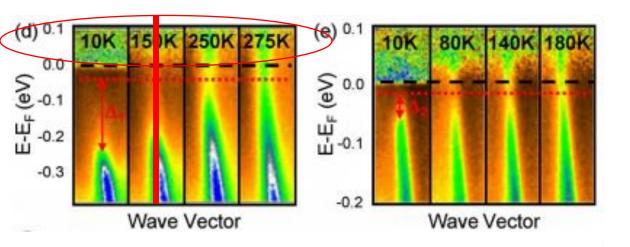
Method 2.5: Energy distribution curve (EDC) + reasonable definition of gap edge

Moore et al. PRB 81, 073102 (2010)

#### Resources

- Campuzano, Norman, Randeria. Photoemission in the high-Tc superconductors. https://arxiv.org/pdf/condmat/0209476.pdf
- Damascelli, Hussain, Shen. Angle-resolved photoemission studies of the cuprate superconductors. Rev. Mod. Phys. 75 473 (2003)
- Damascelli. Probing the Electronic Structure of Complex Systems by ARPES. Physica Scripta. Vol. T109, 61–74, 2004
  - (https://www.cuso.ch/fileadmin/physique/document/Damascelli ARPES CUSO 2011 Lecture Notes.pdf)
- Hufner, Photoelectron Spectroscopy, Springer (2003)

### Extracting spectral gaps



1. Get rid of Fermi-Dirac cutoff by dividing by Fermi-Dirac distribution or symmetrizing (only if there is e-h symmetry)

# Extra: imaging of electrons onto entrance slit via electrostatic lens

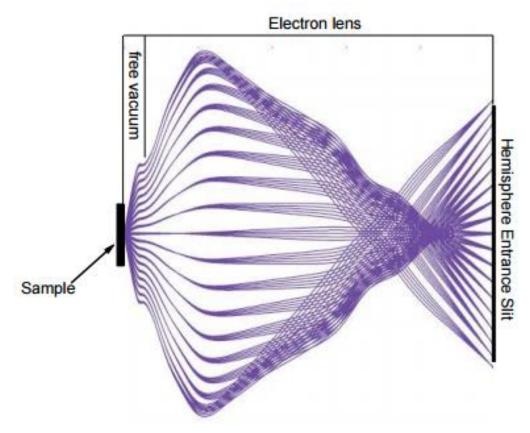


Image from VG Scienta and PhD Thesis of Dr. Ari Deibert Palczewski (http://lib.dr.iastate.edu/cgi/viewcontent.cgi?article=2629&context=etd)